Time reversal and data assimilation

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Statement of the problem



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Let X and Y be Hilbert spaces, $A: \mathcal{D}(A) \to X$ et $C \in \mathcal{L}(X,Y)$.

$$\dot{w}(t) = Aw(t), \ w(0) = x$$
$$y(t) = Cw(t).$$

Assume that A generates a C^0 semigroup, denoted \mathbb{T} , in X and that the pair (A, C) is exactly observable in time τ . Then the map $x \mapsto y$ has a bounded left-inverse.

Practical question: How to compute efficiently x from y?



A naïve solution: inverting the gramian

Solve in the sense of mean squares, i.e., consider the equation

$$Q_{\tau}x = \int_0^{\tau} \mathbb{T}_t^* C^* y(s) \, \mathrm{d}s,$$

where $Q_{\tau} = \int_0^{\tau} \mathbb{T}_t^* C^* C \mathbb{T}_t dt$ is the observability gramian.

Note that $Q_{\tau} > 0$ since (A, C) observable.

The condition number is generally high and the method is very expensive



Outline

• A solution in one shot

• An iterative method

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A solution in one shot



Basic assumptions

 $A = -A^*$, (A, C) exactly observable in time τ and there exists a "time reversal" operator $\mathbf{H}_{\tau} \in \mathcal{L}\left(L^2([0,\tau];X)\right)$ satisfying

$$\mathbf{H}_{\tau}^2 = I,\tag{1a}$$

$$\frac{d}{dt}\mathbf{A}_{\tau}\mathbb{T}_{t}z_{0} = -\mathbf{A}_{\tau}\frac{d}{dt}\mathbb{T}_{t}z_{0}, \qquad \forall z_{0} \in X$$
(1b)

$$A\mathbf{H}_{\tau} \mathbb{T}_t z_0 + \mathbf{H}_{\tau} A \mathbb{T}_t z_0 = 0, \qquad \forall z_0 \in X$$
 (1c)

$$C^*C\mathbf{H}_{\tau}f = \mathbf{H}_{\tau}C^*Cf \qquad \forall f \in C([0,\tau], X), \tag{1d}$$

$$\|(\mathbf{H}_{\tau}f)(0)\| = \|f(\tau)\| \quad \forall f \in C([0,\tau], X).$$
 (1e)

Examples:

 $(\mathbf{H}_{\tau}v)(s) = \overline{v(\tau - s)}$ for the Schrödinger equation.

$$\begin{pmatrix} \mathbf{A}_{\tau} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix} \end{pmatrix} (s) = \begin{bmatrix} v_1(\tau - s) \\ -v_2(\tau - s) \end{bmatrix}$$
 for the wave equation.



A reversed Luenberger observer

$$\dot{v}(t) = (A - C^*C)v(t) + (\mathbf{H}_{\tau}C^*C\mathbf{T}_t x)(t), \qquad t \in (0, \tau)$$

 $v(0) = 0.$

Denote

$$e(t) = v(t) - \mathbf{H}_{\tau} \mathbb{T}_t x.$$

Simple calculations using (1a)-(1e) show that

$$\dot{e}(t) = (A - C^*C)e(t),$$

so that $||e(\tau)|| \le Me^{-\omega \tau} ||e(0)||$.

This means that $v(\tau)$ is a good (and cheap) approximation of x, provided that τ is large.



An iterative method (abstract version of Phung and Zhang, 2008)



The algorithm

For $\gamma > 0$, we define the sequences (v_n) and (e_n) by:

• n = 0:

$$\begin{cases} \dot{v}_0(t) = (A - \gamma C^* C) v_0(t) + \gamma (\mathbf{H}_{\tau} C^* \mathbf{C} \mathbf{T}_t \mathbf{x})(t), & t \in (0, \tau) \\ v_0(0) = 0, \end{cases}$$

$$e_0 = v_0 - \mathbf{A}_{\tau} \mathbb{T}_t x.$$

• $n \ge 1$:

$$\begin{cases} \dot{v}_n(t) = (A - \gamma C^*C)v_n(t) + 2\gamma (\mathbf{A}_{\tau}C^*\mathbf{C}e_{n-1})(t), \\ v_n(0) = 0, \end{cases}$$

$$e_n(t) = v_n(t) - \mathbf{A}_{\tau} e_{n-1}(t).$$



Proposition. The sequences (v_n) and (e_n) satisfy:

- 1. $\dot{e}_n(t) = (A \gamma C^*C)e_n(t), \qquad (n \ge 0).$
- 2. $||e_0(0)|| = ||x||$, $||e_n(0)|| = ||e_{n-1}(\tau)||$ for all $n \ge 1$.
- 3. For every $N \geq 1$ we have :

$$e_{2N}(t) = \left(\sum_{n=0}^{N} v_{2n}(t) - \sum_{n=1}^{N} (\mathbf{H}_{\tau} v_{2n-1})(t)\right) - (\mathbf{H}_{\tau} \mathbb{T}_{t} x)(t). \tag{1}$$

Lemma. Let S be the semigroup generated by $A - \gamma C^*C$ and assume that (A, C) is exactly observable in time τ . Then $||S_{\tau}|| < 1$.



Theorem. Assume that (A, C) is exactly observable in time τ .

$$\left\| x - \sum_{n=0}^{N} (\mathbf{A}_{\tau} v_{2n})(0) \right\| \le \|S_{\tau}\|^{2N} \|x\| \qquad (N \in \mathbb{N}).$$

Proof. By the Lemma we have

$$||e_n(\tau)|| \le ||S_\tau||^n ||e_0(0)|| = ||S_\tau||^n ||x||$$
 $(n \ge 1).$

On the other hand

$$||e_{2N}(\tau)|| = ||(\mathbf{A}_{\tau}e_{2N})(0)|| = \left\| \sum_{n=0}^{N} (\mathbf{A}_{\tau}v_{2n})(0) - x \right\|.$$



Simulations



The wave equation with distributed observation

$$\begin{cases} \ddot{w}(x,t) - \Delta w(x,t) = 0, & x \in \Omega, & t \in (0,\tau), \\ w(x,t) = 0, & x \in \partial\Omega, & t \in (0,\tau), \\ w(x,0) = w_0(x), & x \in \Omega, \\ \dot{w}(x,0) = w_1(x), & x \in \Omega, \end{cases}$$

with the output

$$y = \dot{w}_{|\mathcal{O}},$$

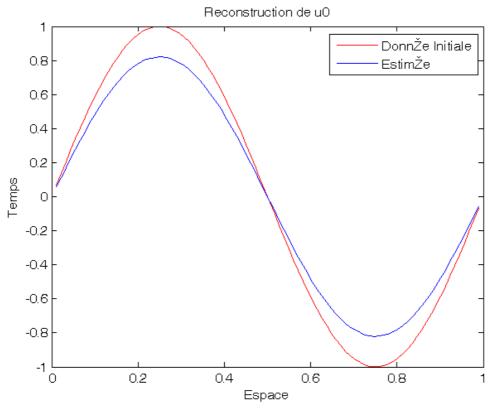
where $\mathcal{O} \subset \Omega$ satisfies the geometric optics condition.



First simulation in one space dimension

$$\Omega = (0,1), \quad \mathcal{O} = \left(\frac{1}{3}, \frac{2}{3}\right), \quad \tau = 2, \quad \gamma = 1,$$

$$w_0(x) = \sin(2\pi x), \quad w_1(x) = 0.$$

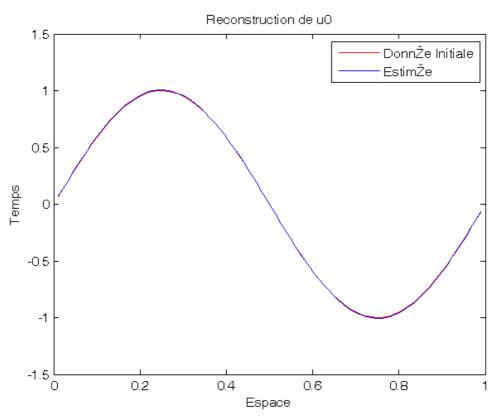




Second simulation in one space dimension

$$\Omega = (0, 1), \quad \mathcal{O} = \left(\frac{1}{3}, \frac{2}{3}\right), \quad \tau = 1, \quad \gamma = 10,$$

$$w_0(x) = \sin(2\pi x), \quad w_1(x) = 0.$$





Extensions and comments



Extensions and comments

- The boundedness of *C* is clearly not necessary, but some weaker assumption on the observation operator are necessary
- Exact observability is not necessary for the convergence
- The choice of the damping coefficient plays a crucial role
- The method can be coupled to a solver of Volterra equations to solve inverse source problems (see Alvez, Silvestre, Takahashi and M.T., 2008)
- It seems possible to generalize the method for operators A with spectrum in a vertical strip
- For a self-adjoint A the problem is strongly ill-posed. Similar algorithms could be interesting to compute the final state.

